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# Molecular open shell configuration interaction calculations using the Dirac-Coulomb Hamiltonian: The $f^6$ -manifold of an embedded $\text{EuO}_6^{9-}$ cluster

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We present results of all-electron molecular relativistic (Hartree-Fock-Dirac) and nonrelativistic (Hartree-Fock) calculations followed by a complete open shell configuration interaction (COSCI) calculation on an  $\text{EuO}_6^{9-}$  cluster in a  $\text{Ba}_2\text{GdNbO}_6$  crystal. The results include the calculated energies of a number of states derived from the  $f^6$ -manifold and  $^5D-^7F$  luminescence transition wavelengths. The calculations were performed using the molecular Fock-Dirac (MOLFDIR) program package developed in our laboratory. The theory and methods employed in this package are briefly described. The physical models used to analyze the  $\text{Eu}^{3+}$  impurity states range from a bare  $\text{Eu}^{3+}$  ion to an  $\text{EuO}_6^{9-}$  cluster embedded in a Madelung potential representing the rest of the crystal. We show that it is necessary to use the embedded cluster model to get a reasonable description of the crystal field splittings of the states arising from the  $f^6$ -manifold. Our results indicate that the calculated splittings are very sensitive to the orbitals used. It is therefore essential that relativistic orbitals be used from the outset.

## I. INTRODUCTION

Compounds containing lanthanide ions ( $Z = 57-71$ ) form an interesting class of materials, both from a scientific and a practical point of view. Their optical and magnetic properties for example are at present technically widely exploited, while being subjects of continuing experimental and theoretical investigations. These properties arise from the complex manifold of electronic states arising from the incompletely filled  $4f$  shells of the lanthanide atoms or ions. Since the classic work of Becquerel and Bethe<sup>1,2</sup> it is well known how to describe these states in crystalline surroundings semiempirically. In recent years this theory has been revitalized and has considerably been extended by Thole in connection with the rationalization of x-ray absorption spectroscopic data.<sup>3,4</sup> At least for the lower lying electronic states, the Coulombic interaction between the electrons (on the order of eVs) still dominates the spin-orbit interaction (on the order of tenths of eVs), although not as heavily as in the case of the  $3d$  metals. In contrast to the latter case, however, the perturbing effects of the surrounding are much smaller than the spin-orbit interaction and therefore smaller than relativistic effects in general. This poses a serious problem when one wants to go beyond a semiempirical description and use *ab initio* computational methods, for example, to obtain more quantitative insight into the origin of crystal fields or into the rates of various radiative or nonradiative decay or excitation transfer processes. The problem is whether *ab initio* results based on the nonrelativistic Schrödinger equation with relativistic corrections added on can be relied upon or that a relativistic approach is needed from the outset to account for the differences between relativistic and nonrelativistic orbitals.

At present the *ab initio* treatment of relativistic effects in compounds of the kind considered here is limited to the use

of an effective Schrödinger equation for valence electrons containing relativistic perturbation terms and a relativistic core potential derived from atomic Fock-Dirac calculations (Ref. 5 and references therein). As it is based on a Schrödinger equation this useful approach allows the inclusion of electron correlation effects by standard quantum chemical methods. All electron *ab initio* calculations, however, including relativistic effects as well as correlation effects on the same theoretical level, have not yet been performed on molecular systems containing heavy atoms. Some calculations based on relativistic quantum mechanics have been performed (Refs. 6-8 and references therein). Most of the calculations concern atoms, and a few deal with diatomic molecules. Applications to molecular systems containing more than two nuclei have been sparse,<sup>9-12</sup> except for calculations based on a local density approach.

More than a decade ago van Piggelen<sup>13</sup> has carried out nonrelativistic *ab initio* calculations, with spin-orbit perturbation corrections, to describe some of the states from the  $f$ -manifold of the  $\text{Eu}^{3+}$  impurity in  $\text{Ba}_2\text{GdNbO}_6$ , using a molecular model (an  $\text{EuO}_6^{9-}$  cluster embedded in the Madelung field of the rest of the crystal). The impurity site in this compound has cubic (octahedral) symmetry, which was important to make the calculations feasible. His interpretation of the calculated crystal field splitting was inconclusive because of the problem mentioned. In this work we use the same molecular model but carry out fully relativistic calculations by applying the all-electron Hartree-Fock-Dirac method followed by relativistic configuration interaction (CI), resulting in a complete intermediate coupling description of the  $f$ -like manifold of the lanthanide impurity. One aim of the work is to obtain results that can serve as a reference for more approximate methods. The method also enables the study of correlation effects and the effects of the Breit interaction.<sup>11</sup>

## II. THEORY

We use an average of configurations self-consistent-field (SCF) method (Hartree–Fock for nonrelativistic calculations, Hartree–Fock–Dirac for relativistic calculations) to generate a set of one-electron spin orbitals. Using these average orbitals, we perform a complete open shell configuration interaction (COSCI) calculation to get a description of the open shell manifold. In this section we give the basic equations of the method and the way these are solved. Some details of the molecular Fock–Dirac (MOLFDIR) program package<sup>9,11,14,15</sup> are given, in particular the form of the basis functions, kinetic and atomic balance, and general contraction.

### A. General

The time-independent Dirac equation for one-electron molecular systems<sup>16,17</sup> is given by

$$h\varphi = \varepsilon\varphi, \quad (1)$$

in which  $h$  is the one-electron Dirac operator defined by

$$h = c\boldsymbol{\alpha} \cdot \mathbf{p} + (\beta - 1)mc^2 + V. \quad (2)$$

We have shifted the energy scale by  $-mc^2$  to facilitate comparison with nonrelativistic energies. The  $\alpha$  and  $\beta$  are  $4 \times 4$  matrices defined by

$$\alpha = \begin{pmatrix} 0 & \sigma \\ \sigma & 0 \end{pmatrix}$$

and

$$\beta = \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix} \quad (3)$$

in which  $\sigma$  is the collection of Pauli spin-matrices,  $V$  is the potential energy  $V(\mathbf{r}; \mathbf{R})$  due to the nuclei at  $\mathbf{R}$ ,  $c$  is the speed of light in vacuum (137.036 02 a.u. used), and  $m$  is the rest-mass of the electron.  $\varphi(\mathbf{r})$  is a four-component spinor, conveniently written in bispinor form,

$$\varphi = \begin{pmatrix} \varphi^L \\ \varphi^S \end{pmatrix}. \quad (4)$$

The first-order approximation to the positive-energy state is

$$\varphi^S \approx \frac{1}{2mc} \boldsymbol{\sigma} \cdot \mathbf{p} \varphi^L \quad (5)$$

so the lower bispinor is conventionally called the “small component,” and the upper bispinor is the “large component.”

One can generalize the one-electron Dirac equation to an approximate relativistic many-electron equation (the Dirac–Coulomb equation),<sup>18</sup> which is given (in atomic units) by

$$\left\{ \sum_i [\boldsymbol{\alpha}_i \cdot \mathbf{p}_i + (\beta_i - 1)mc^2 + V_i] + \sum_{i < j} \frac{1}{r_{ij}} \right\} \psi = E\psi. \quad (6)$$

The two-electron interaction in this equation consists of the usual Coulomb term and is not relativistically invariant; the leading correction to the two-electron interaction (the Breit interaction<sup>19</sup>) can be included either variationally or by perturbation theory,<sup>9,20,21</sup> but has been left out in this work.

### B. Open shell approach

The molecules or clusters, modeling bulk or surface properties of solids, we want to study contain heavy atoms and are in general open shell systems. Especially lanthanides and actinides have an open  $f$ -shell, which gives rise to a large number of energy eigenstates lying close together. A simple open shell SCF method alone is not sufficient to describe such a manifold. First, significant interaction between the Russell–Saunders terms of the  $f$ -multiplet should be expected. Second, since (as is generally known) the correct eigenfunctions of the Dirac–Coulomb Hamiltonian are neither pure  $LS$ -coupled functions nor pure  $JJ$ -coupled functions, a many-electron function in intermediate coupling should be constructed. This can be accomplished by forming linear combinations of determinants constructed from  $LS$  or  $JJ$ -coupled one-electron spinors. Furthermore, since we are interested in the optical spectrum of the  $f$ -manifold, we need to calculate a large number of the energy eigenstates. A separate SCF calculation for each of these states is cumbersome, and is certainly not a practical approach.

By using the COSCI approach these problems can be solved. We start with a SCF calculation using an average energy expression [defined as the average of the energy expressions for all possible individual states of the relevant configuration(s)], resulting in the “average” total energy and a set of “average” orbitals. Next, a CI calculation is performed within the full-CI space generated from the set of open shell orbitals from the SCF calculation. This results in a description of the open shell manifold which is not tied to one of the extreme coupling schemes.

### C. The relativistic open shell SCF equations

The relativistic open shell SCF equations are formally derived in the same way as the nonrelativistic equations.<sup>22</sup> The one-electron Schrödinger operators are replaced by Dirac operators and the scalar orbitals are now four component spinors. Fundamentally there is of course an important difference in the meaning of the variational method since it is used to locate a stationary state in the positive energy domain instead of an absolute minimum. The resulting SCF equations are called the Hartree–Fock–Dirac equations.

The average of configuration energy expression which is used in the MOLFDIR program package has the following form:<sup>14</sup>

$$E = \sum_k h_k + \frac{1}{2} \sum_{kl} Q_{kl} + f \left[ \sum_m^o h_m + \frac{1}{2} a f \sum_{mn}^o Q_{mn} + \sum_{k,m}^{c,o} Q_{km} \right]. \quad (7)$$

In this equation,  $k$  and  $l$  label closed shell spin-orbitals (which, by definition, are occupied by one electron),  $m$  and  $n$  label open shell spin-orbitals (with related to them a fractional occupation number  $f$  and a coupling constant  $a$ ),  $h_k$  is the diagonal matrix element of the one-electron Dirac operator [from Eq. (1)] over spin-orbital  $\varphi_k$ , and  $Q_{ij} \equiv J_{ij} - K_{ij}$  in which  $J_{ij}$  and  $K_{ij}$  are the usual Coulomb and exchange integrals. Apart from the use of four-component spin-orbitals instead of orbitals, the energy expression differs from the energy expression used by Roothaan in the detailed form of the one-electron operators (Dirac-operators vs one-electron

Schrödinger operators) and of the two-electron operators when the Breit interaction is included.

If we use the average of configuration total energy expression, the coupling constant and fractional occupation number are given as functions of the number of open shell electrons  $n$  and the number of open shell orbitals  $d$ ,

$$f = \frac{n}{d}, \quad (8)$$

$$a = \frac{d(n-1)}{n(d-1)}. \quad (9)$$

For completeness, we present the working formulas which are obtained when the SCF equations are solved using a basis set (consisting of either two or four component spinors) expansion technique. In these equations  $p, q, r, \dots$  label basis spinors  $\chi_p, \chi_q, \chi_r, \dots$  and  $S_{pq}$  is the overlap between basis spinors  $\chi_p$  and  $\chi_q$ , and  $(pq||rs)$  is an antisymmetrized two-electron repulsion integral in charge-cloud notation,

$$\alpha = \frac{1-a}{1-f}, \quad (10)$$

$$|i\rangle = \sum_p \chi_p c_{pi}, \quad (11)$$

$$D_{pq}^C = \sum_k c_{qk}^* c_{pk}, \quad (12)$$

$$D_{pq}^O = f \sum_m c_{qm}^* c_{pm},$$

$$Q_{pq}^C = \sum_{rs} (pq||rs) D_{sr}^C, \quad (13)$$

$$Q_{pq}^O = \sum_{rs} (pq||rs) D_{sr}^O,$$

$$L_{pq}^C = \sum_{rs} [S_{pr} D_{rs}^C Q_{sq}^O + Q_{pr}^O D_{rs}^C S_{sq}], \quad (14)$$

$$L_{pq}^O = \sum_{rs} [S_{pr} D_{rs}^O Q_{sq}^O + Q_{pr}^O D_{rs}^O S_{sq}],$$

$$F^C = h + Q^C + Q^O + \alpha L^O, \quad (15)$$

$$F^O = h + Q^C + a Q^O + \alpha L^C.$$

For the closed shell orbitals we get the equation

$$F^C |k\rangle = \varepsilon_k |k\rangle, \quad (16)$$

and for the open shell orbitals we get the equation

$$F^O |m\rangle = \varepsilon_m |m\rangle. \quad (17)$$

Equations (16) and (17) need to be solved iteratively. The coupling operators  $L^C$  and  $L^O$  take care of the orthogonality between the closed shell and open shell solutions. We get the following expression for the total energy:

$$E = \text{Tr}[H^D D^C] + \frac{1}{2} \text{Tr}[Q^C D^C] + \text{Tr}[H^D D^O] + \frac{1}{2} a \text{Tr}[Q^O D^O] + \text{Tr}[Q^O D^C]. \quad (18)$$

## D. The COSCI method

In this section, we mention some features of the COSCI method. The many-electron function space used in the COSCI calculations is the full-CI space generated from a (small) set of active spin-orbitals which is usually chosen to

be the set of open shell spin-orbitals from the SCF calculation. Since the average energy is proportional to the trace of the Hamiltonian matrix in this many-electron basis, it is invariant for unitary transformations of the one-electron basis. In particular, it does not make any difference if we had started with an  $LS$  coupled instead of a  $JJ$  coupled basis set.

Since only the open shell space is used to construct the CI matrix, correlation corrections are hardly accounted for at this stage. Also, the open shell orbitals are not optimized for individual states. The method is expected, however, to yield a balanced, global picture of the open shell manifold of states. More accurate descriptions of specific states can be found subsequently by the application of multireference configuration interaction (MRCI) or multiconfiguration self-consistent field (MCSCF) techniques.

We define the set of frozen orbitals to consist of all one-electron spinors which are present in all determinants spanning the CI-space. The interaction with these frozen orbitals is taken into account by using an effective one-electron operator  $h'$ ,

$$h' = h + \sum_f^F (J_f - K_f), \quad (19)$$

where  $J_f$  is the Coulomb operator with spin-orbital  $\varphi_f$ ,  $K_f$  is the exchange operator with spin-orbital  $\varphi_f$ , and the summation runs over the set of frozen spin-orbitals  $F$ . The matrix elements of this operator can be calculated easily using the fact that the one-electron spinors are solutions of the SCF Eqs. (16) and (17).

Finally, it is clear that while the COSCI method has been introduced here to perform open shell calculations with a relativistic Hamiltonian, it can also be used for calculations with a nonrelativistic Hamiltonian.

## E. Basis functions

The MOLFDIR package works with two distinct sets of, usually atom centered, scalar (contracted) Cartesian Gaussian functions, a large component set  $\{g_i^L\}$  and a small component set  $\{g_i^S\}$ . The two-electron repulsion integrals are calculated over the functions belonging to these sets. From these two scalar sets, two new sets of symmetry adapted molecular basis spinors  $\{\chi_i^L\}$  and  $\{\chi_i^S\}$  are constructed using the Dirac double group symmetry elements. The spinors  $\chi_i^L$  and  $\chi_i^S$  are defined by

$$\chi_i^L = \begin{pmatrix} \chi_i^{L\alpha} \\ \chi_i^{L\beta} \\ 0 \\ 0 \end{pmatrix} \quad \text{and} \quad \chi_i^S = \begin{pmatrix} 0 \\ 0 \\ \chi_i^{S\alpha} \\ \chi_i^{S\beta} \end{pmatrix}, \quad (20)$$

$$\chi_i^{L\alpha} = \sum_j g_j^L c_{ij}^{L\alpha}, \quad \chi_i^{L\beta} = \sum_j g_j^L c_{ij}^{L\beta},$$

$$\chi_i^{S\alpha} = \sum_j g_j^S c_{ij}^{S\alpha}$$

and

$$\chi_i^{\sigma\beta} = \sum_j g_{ij}^{\sigma} c_{is}^{\sigma\beta}, \quad (21)$$

where the coefficients  $c_{ij}^{\sigma}(X \in \{L, S\}; \sigma \in \{\alpha, \beta\})$  in Eq. (21) are determined by symmetry. Thus, each nonzero component of these basis spinors consists of a linear combination of the basis functions from set  $\{g_i^{\sigma}\}$  or set  $\{g_i^{\beta}\}$ .

#### F. Kinetic balance and general contraction

As has been discussed by several authors,<sup>23–25</sup> it is important to ensure that the basis set in which the one-electron functions are expanded fulfills the kinetic balance condition. This condition requires that for each function  $\chi_i^{\sigma}$  in the large component basis, the function  $\alpha \cdot \mathbf{p} \chi_i^{\sigma}$  is contained in the small component basis set. Since kinetic balance is neither a necessary nor sufficient condition to produce the best approximate eigensolutions, we use an extended kinetic balance scheme.<sup>15</sup> In practice this means that the small component basis set contains more basis functions than those minimally required by the kinetic balance condition.

The first set of extra functions results from our choice to employ two separate primitive Gaussians instead of the fixed linear combination that results from  $\alpha \cdot \mathbf{p}$  operating on a single Gaussian primitive in the large component.

A second set of extra functions is used in molecular calculations where atomic solutions are used to contract the large and small component basis functions. In addition to the kinetically balanced functions the atomic small component solutions are used as small component basis functions. A basis set which includes such an extension is called an atomically balanced basis set.<sup>15</sup>

When general contraction<sup>26</sup> is used to reduce the variational space we still require that the small component basis fulfills atomic and kinetic balance conditions. The contraction reduction is thereby counteracted by the fact that up to four contracted functions for each large component function are necessary (up to two functions to ensure kinetic balance, and up to two functions to ensure atomic balance). Although these small component functions may tend to form a linear dependent set, in which case one or more functions can be removed, the gain obtained by the use of general contraction in the small component basis set is small compared to that obtained in the large component basis set.

### III. The Ba<sub>2</sub>GdNbO<sub>6</sub>:Eu system

#### A. Description of the system

In 1966, Blasse *et al.*<sup>27,28</sup> have obtained experimental data (luminescence spectra) on some compounds which contained an Eu<sup>3+</sup> impurity. Due to the approximate validity of spin and parity selection rules, electric dipole transitions between the <sup>5</sup>D and <sup>7</sup>F levels of the Eu<sup>3+</sup> impurity are forbidden in first approximation. The spin selection rule is lifted by the spin–orbit coupling. The magnetic dipole and quadrupole transitions are only forbidden by a spin selection rule, which is again lifted by the spin–orbit coupling. To investigate the effect of the symmetry of the surroundings, a series of compounds was studied, a compound in which the Eu<sup>3+</sup> ion occupies a strict center of (*O<sub>h</sub>*) symmetry, a compound in which small deviations from this symmetry occur,

and a compound with no symmetry at all. In the last case, both electric and magnetic dipole transitions can occur, while in the first case only magnetic dipole transitions are allowed.<sup>29,30</sup>

We have chosen to study a compound with a strict center of symmetry, Ba<sub>2</sub>GdNbO<sub>6</sub>:Eu. The advantage is that calculations can be performed using *O<sub>h</sub>* symmetry, which greatly reduces the computational efforts needed (both in terms of CPU time and of disk space).

The Ba<sub>2</sub>GdNbO<sub>6</sub> crystal (Fig. 1) has an ordered perovskite structure.<sup>31</sup> The unit cell of the crystal is a cube divided in eight smaller cubes. In the center of the smaller cubes we find alternating Gd<sup>3+</sup> and Nb<sup>5+</sup> ions, on the corners of the smaller cubes we find Ba<sup>2+</sup> ions, and at the middle of the sides of the small cubes the O<sup>2−</sup> ions are positioned. The Eu<sup>3+</sup> impurity occupies one of the Gd<sup>3+</sup> sites with symmetry *O<sub>h</sub>*. It is surrounded by six O<sup>2−</sup> ions, forming an “EuO<sub>6</sub><sup>9−</sup> cluster,” with an Eu–O distance of 4.02 a.u.

The experimental results obtained by Blasse *et al.* (the energy distribution of the emission of Eu<sup>3+</sup>-activated Ga<sub>2</sub>GdNbO<sub>6</sub>) show a sharp line at 595 nm (16 800 cm<sup>−1</sup>) corresponding to the <sup>5</sup>D<sub>0</sub> → <sup>7</sup>F<sub>1</sub> transition. This is an allowed magnetic dipole transition. Also, at least six weak and broad lines are visible in the region 610–630 nm (15 900–16 400 cm<sup>−1</sup>). These lines were assigned to the <sup>5</sup>D<sub>0</sub> → <sup>7</sup>F<sub>2</sub> transitions (the large number of lines is due to simultaneous vibronic transitions).

#### B. Physical model

We have performed all-electron average of configuration open shell Hartree–Fock and Hartree–Fock–Dirac calculations, followed by COSCI calculations [using, respectively, the Schrödinger Hamiltonian (*NR*) and the Dirac–Coulomb Hamiltonian (*FD*)], on part of the Ba<sub>2</sub>GdNbO<sub>6</sub> crystal together with the impurity. This part of the material, cut from the complete system, will be referred to as “the cluster,” even when “the cluster” consists of one atom only (the impurity atom). The influence of the surrounding infinite crystal is accounted for by considering the electrostatic effects on the cluster only. The Madelung potential of the

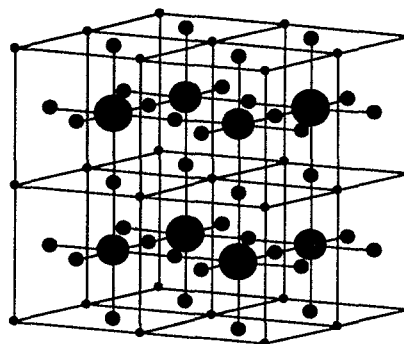


FIG. 1. The Ba<sub>2</sub>GdNbO<sub>6</sub> unit cell. In order to get a better view of the structure of the cell, the ions have not been drawn to scale. Small black spheres, Ba<sup>2+</sup>; small grey spheres O<sup>2−</sup>, remaining large spheres, alternating Gd<sup>3+</sup> and Nb<sup>5+</sup>. The Eu<sup>3+</sup> impurity occupies a Gd<sup>3+</sup> site.

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delung potential in the cluster region has been calculated. After subtracting the cluster contributions, the remaining potential was fitted with a set of point charges. The fit results for the two different clusters (the  $\text{Eu}^{3+}$  ion, and the  $\text{EuO}_6^{9-}$  system) are given in Table III.

For the interpretation of the results it is useful to note that the Madelung potential used with the  $\text{Eu}^{3+}$  ion can be fitted with only six charges on the  $\text{O}^{2-}$  sites (at a distance of 4.02 a.u.). The fit charges are  $-0.865$ , with a standard deviation in the potential of 0.04 a.u.<sup>13</sup> In the embedded ion calculations, however, we have used the much more accurate fit given in Table III.

### C. The $\text{EuO}_6^{9-}$ cluster

In the Hartree–Fock–Dirac calculations on the  $\text{EuO}_6^{9-}$  clusters, using the experimental geometry, the symmetry unique two-electron integrals which were larger than  $10^{-12}$  for the  $(LL|LL)$  and  $(LL|SS)$  integrals, and the  $10^{-8}$  for the  $(SS|SS)$  integrals were calculated. This resulted in  $2 \times 10^6$   $(LL|LL)$  integrals,  $48 \times 10^6$   $(LL|SS)$  integrals and  $241 \times 10^6$   $(SS|SS)$  integrals. If no symmetry would have been employed, we would have obtained  $0.1 \times 10^9$   $(LL|LL)$  integrals,  $2.2 \times 10^9$   $(LL|SS)$  integrals, and  $12.7 \times 10^9$   $(SS|SS)$  integrals.

We found that the convergence of the SCF iterations was very slow. Normally, the computational efforts needed for a single SCF iteration can be reduced by temporarily using only part of the large-component two-electron integrals, or by temporarily fixing the contributions of the small component two-electron integrals to the Fock matrices; in this case we found that both of these methods did seriously degrade the convergence behavior.

We also found that selection of the open shell eigenvectors by overlap with the eigenvectors of the previous iteration did not work very well for the  $\text{EuO}_6^{9-}$  systems. However, selection by character (thus forcing the open shell

orbitals to be constructed from the Eu  $f$ -basis spinors) worked much better.

## V. RESULTS

### A. The free $\text{Eu}^{3+}$ ion

In order to get an estimate of the accuracy which can be obtained using the given general contracted basis set together with the COSCI approach, we have performed some calculations on the  $\text{Eu}^{3+}$  ion in the nonrelativistic limit: (1) Numerical calculations using the NMCHF program of Froese-Fischer,<sup>35</sup> these calculations are used as reference. (2) Atomic basis set expansion calculations with the ASCF (Ref. 13) program, using the uncontracted basis set described before. (3) MOLFDIR–COSCI calculations using the general contracted basis set. The results are given in Table IV. All absolute energies are too high by  $\sim 0.034$  a.u., due to the finite basis set approximation. The relative errors show the accuracy within the  $f$ -manifold, which is in general much better than  $10^{-3}$  a.u. For the  $^7F$  and  $^5L$ -states, the error due to the use of a set of average orbitals (column III – column II in Table IV) is significantly larger than for the other states. This error, especially in case of the  $^5L$ -state, is partially canceled by the basis set error (column II – column I in Table IV).

We conclude that in the worst case the accuracy of the calculations, as far as splittings between Russell–Saunders term are concerned, can be as bad as  $3 \times 10^{-3}$  a.u. ( $\sim 600 \text{ cm}^{-1}$ ), and that this is mainly due to the average  $f$ -shell approach. The accuracy in the many-electron total energies will therefore not significantly be improved by using a more flexible basis set. The splittings within Russell–Saunders terms due to relativistic and ligand field effects are much smaller, and the absolute accuracy of these splittings is expected to be much better.

TABLE III. Fit to the Madelung potential for several clusters. Position = position of the fit-charges in units of 4.02 a.u.; points = number of symmetry equivalent points for that site. All numbers are in a.u.

	Madelung	Fit for $\text{EuO}_6$	Fit for Eu
$V_{\text{Eu}}^a$	− 1.322 127 3	1.662 956 <sup>c</sup>	− 1.3222 084 2
$(\Delta V)_{\text{max}}^b$		$3.5 \times 10^{-4}$	$5.0 \times 10^{-5}$
$(\Delta V)_{\text{av}}^c$		$2.1 \times 10^{-7}$	− $2.4 \times 10^{-5}$
$\sigma(V)^d$		$6.7 \times 10^{-5}$	$2.5 \times 10^{-5}$

Site	Position	points	Fit charge for $\text{EuO}_6$ cluster	Fit charge for Eu cluster
$\text{O}^{2-}$	(0,0,1)	6		− 1.998 65
$\text{Ba}^{2+}$	(1,1,1)	8	1.962 66	1.996 09
$\text{Nb}^{5+}$	(0,0,2)	6	4.970 18	4.640 23
$\text{O}^{2-}$	(2,0,1)	24	− 1.915 79	− 1.533 86
$\text{Gd}^{3+}$	(2,2,0)	12	1.590 02	
$\text{Ba}^{2+}$	(3,1,1)	24	− 0.480 09	

<sup>a</sup> Potential at  $\text{Eu}^{3+}$  site.

<sup>b</sup> Largest difference between the fitted potential and the exact potential.

<sup>c</sup> Average difference between the fitted potential and the exact potential.

<sup>d</sup> Standard deviation in fitted potential.

<sup>e</sup> – 1.322 118 5 when the contribution of six  $\text{O}^{2-}$  ions at 4.02 a.u. is included.



TABLE IV. Comparison of numerical (NMCHF), basis set (ASCF), and COSCI total energies relative to the average total energy (in a.u.).

State	NMCHF <sup>a</sup> (I)	ASCF <sup>b</sup> (II)	COSCI <sup>c</sup> (III)	II – I	III – II
<sup>7</sup> F	– 0.358 73	– 0.358 83	– 0.355 93	– 0.000 10	0.002 90
<sup>5</sup> L	– 0.225 42	– 0.225 83	– 0.224 73	– 0.000 42	0.001 10
<sup>5</sup> K	– 0.150 29	– 0.150 66	– 0.150 14	– 0.000 37	0.000 52
<sup>3</sup> O	– 0.138 84	– 0.139 10	– 0.138 72	– 0.000 26	0.000 38
<sup>3</sup> N	– 0.111 07	– 0.111 47	– 0.111 23	– 0.000 41	0.000 24
<sup>5</sup> P	– 0.110 43	– 0.110 57	– 0.110 28	– 0.000 15	0.000 29
<sup>5</sup> S	– 0.047 81	– 0.047 50	– 0.047 46	0.000 31	0.000 05

<sup>a</sup>  $E_{av}$  (NMCHF) = – 10 421.679 15.<sup>b</sup>  $E_{av}$  (ASCF) = – 10 421.644 89.<sup>c</sup>  $E_{av}$  (COSCI) = – 10 421.644 89.

## B. The Eu<sup>3+</sup> impurity

From the average of configuration open shell SCF calculations we have obtained the *f* or *f*-like orbital energies, which are given in Tables V and VI. In Fig. 2 the splitting of these orbital energies due to relativistic and crystal field effects is pictorially represented. In the figure the *FD* orbital energies have been shifted upward by 0.18 a.u. The additional shift (Tables V and VI) in the embedded ion results corresponds closely to the value of the Madelung potential at the Eu<sup>3+</sup> site. To explain the shift seen in the embedded cluster results also the interaction with the rest of the cluster needs to be taken into account. We find that this interaction can be fitted (using data from Tables III, V, and VI) with six effective charges of – 1.5 at the O<sup>2–</sup> ion positions, or with six charges of – 2.0 at a distance of 5.3 a.u.

The COSCI results concerning the <sup>7</sup>F and the lower of the <sup>5</sup>D levels are presented in Tables VII and VIII. In Fig. 3 the *FD* data on the <sup>7</sup>F levels is presented in the form of density of states plots. These plots have been produced by using a convolution of a Gaussian with the discrete COSCI eigenvalues. The width of the Gaussian was chosen to match the accuracy of the calculations. In Fig. 4 the EuO<sub>6</sub>MP *FD* COSCI results for the complete *f*-manifold are shown.

Going from the free ion to the embedded ion, certain total energy levels (depending on the symmetry) and the 4*f* orbital energies are split due to the Madelung field representing the crystal. When we include the six O<sup>2–</sup> ions in the cluster together with the Madelung field of the rest of the crystal, we find that the total energy splittings are reduced significantly with respect to the embedded ion results (Fig. 3, Tables VII and VIII). This behavior, which is opposite to the behavior of the transition metal ions,<sup>36</sup> can be explained

by the presence of the filled Eu<sup>3+</sup> 5*s* and 5*p* shells. The superimposed charge distributions of these shells and the oxygen ions becomes slightly deformed. A small amount of charge is shifted from the Eu–O bond region into the regions between these bonds. This in effect reduces the electrostatic splitting of the 4*f* orbitals. Also, due to the localized nature of the 4*f* orbitals, covalent effects are not as important as in the case of transition metal ions, where they lead to an increase of the splitting.

In the nonrelativistic case, crystal field theory based on free ion orbitals yields the same splittings respectively for the many-electron <sup>7</sup>F level and the 4*f* orbital energy. Our embedded ion results reflect this behavior rather closely, in contrast to the embedded cluster results that show substantial deviations. This is due to changes in the detailed form of the 4*f*-like spin-orbitals. From this sensitivity it can be concluded that it is important to use “relativistic” orbitals, which deviate significantly from nonrelativistic orbitals, from the outset.

The much larger splittings calculated for the bare EuO<sub>6</sub><sup>9–</sup> cluster should be compared with the splittings which would result from a calculation on the ion with O<sup>2–</sup> ions represented by six two-point charges, not including the rest of the crystal. These splittings can be estimated from our embedded ion results by assuming that the splittings are linear in the charge of the point charges. The complete Madelung potential can approximately be fitted by six charges of – 0.865 at the O<sup>2–</sup> positions,<sup>13</sup> so we should compare the splittings calculated using the bare cluster with the splittings calculated using the embedded ion multiplied by 2.3

TABLE V. Relative and average *NR f*-electron orbital energies (a.u.)

Level	Eu <i>NR</i>	Eu <i>NR</i>	EuO <sub>6</sub> MP <i>NR</i>	EuO <sub>6</sub> <i>NR</i>
<i>a</i> <sub>2u</sub>	0.000 00	0.000 00	0.000 00	0.000 00
<i>t</i> <sub>2u</sub>	0.000 00	0.002 10	0.002 58	0.002 54
<i>t</i> <sub>1u</sub>	0.000 00	0.004 20	0.005 00	0.007 33
Average	– 1.778 87	– 0.458 02	– 1.142 73	0.748 40

TABLE VI. Relative and average *FD f*-electron orbital energies (a.u.).

Level	Eu <i>FD</i>	EuMP <i>FD</i>	EuO <sub>6</sub> MP <i>FD</i>	EuO <sub>6</sub> <i>FD</i>
<i>e</i> <sub>2u</sub>	0.000 00	0.000 00	0.000 00	0.000 00
<i>f</i> <sub>u</sub>	0.000 00	0.002 62	0.002 13	0.003 46
<i>e</i> <sub>2u</sub>	0.024 10	0.023 80	0.023 80	0.023 69
<i>f</i> <sub>u</sub>	0.024 10	0.026 16	0.025 63	0.026 54
<i>e</i> <sub>1u</sub>	0.024 10	0.027 71	0.026 95	0.028 96
Average	– 1.595 26	– 0.274 29	– 0.970 88	0.923 39

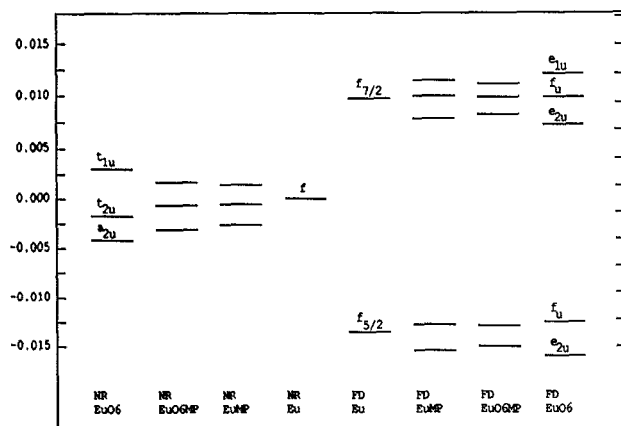


FIG. 2. Calculated splittings (in a.u.) of the orbital energies of the  $f$  and  $f$ -like orbitals due to relativistic and crystal field effects, relative to the average energies.

( $= 2/0.865$ ). Now we find a similar trend as before, inclusion of the six  $O^{2-}$  ions reduces the calculated splittings.

The differences going from embedded ion to embedded cluster are relatively large (of the same order of magnitude as the crystal field splittings), and one might ask whether the cluster should be enlarged. The distance from the  $\text{Eu}^{3+}$  ion to the next-nearest-neighbors (the  $\text{Ba}^{2+}$  ions) is rather large ( $\sim 7$  a.u.), and inclusion of these ions in the cluster is therefore not expected to alter the results significantly. It is expected, however, that the use of point charges deforms the oxygen charge distributions somewhat.

### C. The spectrum

The dominant luminescence transitions occurring in this system are basically the atomic  ${}^5D_0 \rightarrow {}^7F_1$  and  ${}^5D_0 \rightarrow {}^7F_2$  transitions. The first of these is a magnetic dipole transition which can be seen clearly in the spectrum. The second transition is an induced electric dipole transition, allowed by vibronic coupling to phonon modes that break the inversion symmetry. This transition is split by the reduction of the symmetry. In Table IX the calculated transition wave numbers are given.

Based on the  $\text{EuO}_6\text{MP}$  *FD* results, we expect that the splitting of the  $^5D_0 \rightarrow ^7F_2$  transition is of the order of  $100 \text{ cm}^{-1}$ . The free ion model shows no splitting. When we introduce the Madelung potential the splitting is somewhat overestimated ( $251 \text{ cm}^{-1}$ ). Inclusion of the neighboring  $\text{O}^{2-}$  ions in the model ( $\text{EuO}_6\text{MP}$ ) reduces the splitting to

TABLE VII. *NR* COSCI energies in a.u. of the lowest states of the  $f^6$ -manifold with their degeneracies.

	Eu <i>NR</i>			EuMP <i>NR</i>		EuO <sub>6</sub> MP <i>NR</i>		EuO <sub>6</sub> <i>NR</i>	
<sup>7</sup> <i>F</i>	0.000 00	49	0.000 00	21	0.000 00	21	0.000 00	21	
			0.002 28	21	0.001 38	21	0.003 43	21	
			0.004 52	7	0.003 17	7	0.005 10	7	
<sup>5</sup> <i>D</i>	0.114 78	25	0.116 07	15	0.114 07	15	0.114 72	15	
			0.116 70	10	0.114 35	10	0.115 45	10	

TABLE VIII. *FD* COSCI energies in a.u. of the lowest states of the  $f^6$ -manifold with their degeneracies.

	Eu <i>FD</i>			EuMP <i>FD</i>			EuO <sub>6</sub> MP <i>FD</i>			EuO <sub>6</sub> <i>FD</i>		
<sup>7</sup> F <sub>0</sub>	0.000 00	1		0.000 00	1		0.000 00	1		0.000 00	1	
<sup>7</sup> F <sub>1</sub>	0.001 71	3		0.001 66	3		0.001 70	3		0.001 66	3	
<sup>7</sup> F <sub>2</sub>	0.004 82	5		0.003 99	2		0.004 49	2		0.004 08	2	
				0.005 13	3		0.004 99	3		0.005 07	3	
<sup>7</sup> F <sub>3</sub>	0.008 94	7		0.008 51	1		0.008 71	1		0.008 54	1	
				0.009 17	3		0.008 94	3		0.008 94	3	
				0.009 35	3		0.009 02	3		0.009 21	3	
<sup>7</sup> F <sub>4</sub>	0.013 77	9		0.012 94	3		0.013 36	3		0.012 85	3	
				0.014 20	2		0.013 77	2		0.014 19	2	
				0.014 66	3		0.013 98	3		0.014 30	3	
				0.015 08	1		0.014 24	1		0.014 41	1	
<sup>7</sup> F <sub>5</sub>	0.019 10	11		0.018 66	3		0.018 76	3		0.018 57	3	
				0.018 81	2		0.018 81	2		0.018 74	2	
				0.019 21	3		0.019 16	3		0.018 80	3	
				0.020 09	3		0.019 30	3		0.019 91	3	
<sup>7</sup> F <sub>6</sub>	0.024 74	13		0.024 12	2		0.024 30	2		0.023 98	2	
				0.024 25	3		0.024 33	3		0.024 11	3	
				0.024 54	1		0.024 44	1		0.024 41	1	
				0.026 06	3		0.024 94	3		0.025 42	3	
				0.026 22	3		0.025 03	3		0.025 59	3	
				0.026 37	1		0.025 11	1		0.025 73	1	
<sup>5</sup> D <sub>0</sub>	0.093 70	1		0.093 83	1		0.092 24	1		0.092 18	1	
<sup>5</sup> D <sub>1</sub>	0.101 21	3		0.101 32	3		0.099 72	3		0.099 61	3	
<sup>5</sup> D <sub>2</sub>	0.112 75	5		0.112 63	3		0.111 15	3		0.110 84	3	
				0.113 16	2		0.111 25	2		0.111 28	2	

110 cm<sup>-1</sup>. This number is to be compared with the number found by van Piggelen (139 cm<sup>-1</sup>).<sup>13</sup> Although the calculated splitting is quite consistent with the experimental data, a more precise comparison cannot be made at present. The observed vibronic transitions have not been assigned and in fact more and better resolved data are required to do this.<sup>37</sup> The EuO<sub>6</sub>MP model, however, should yield a reasonable description of the effects of the surroundings. We tend therefore to conclude that a crystal field model based on just the Madelung field overestimates the splitting.

The spin-orbit splitting (using the weighted average of the  $E_g$  and the  $T_{2g}$  levels) is calculated to be 628, 662, and 678  $\text{cm}^{-1}$  for, respectively, Eu, EuMP, and EuO<sub>6</sub>MP. This is in good agreement with the experimental data from which we estimate the splitting to be between 600 and 700  $\text{cm}^{-1}$ . van Piggelen's results yield a somewhat larger value (804  $\text{cm}^{-1}$ ). Except for the neglect of the Breit interaction we have not introduced any approximations in the description of the relativistic effects, so we expect that our results are reliable at this point.

The calculated splitting between the  ${}^5D$  level and the  ${}^7F$  level is much too large, in all calculations. For the  ${}^5D_0 \rightarrow {}^7F_1$  transition, we have a discrepancy with experiment of  $\sim 3000$   $\text{cm}^{-1}$ . This large error cannot be attributed to defects of the COSCI or basis set approach (in Sec. V A we have seen that the worst of these errors is of order  $600$   $\text{cm}^{-1}$ ). This difference is of course to be expected at the present level of theory since most of the correlation effects have not yet been included.

## VI. CONCLUSIONS

In this paper we have given results of calculations on  $\text{EuO}_9^{9-}$  clusters in a Madelung potential, as a model for an

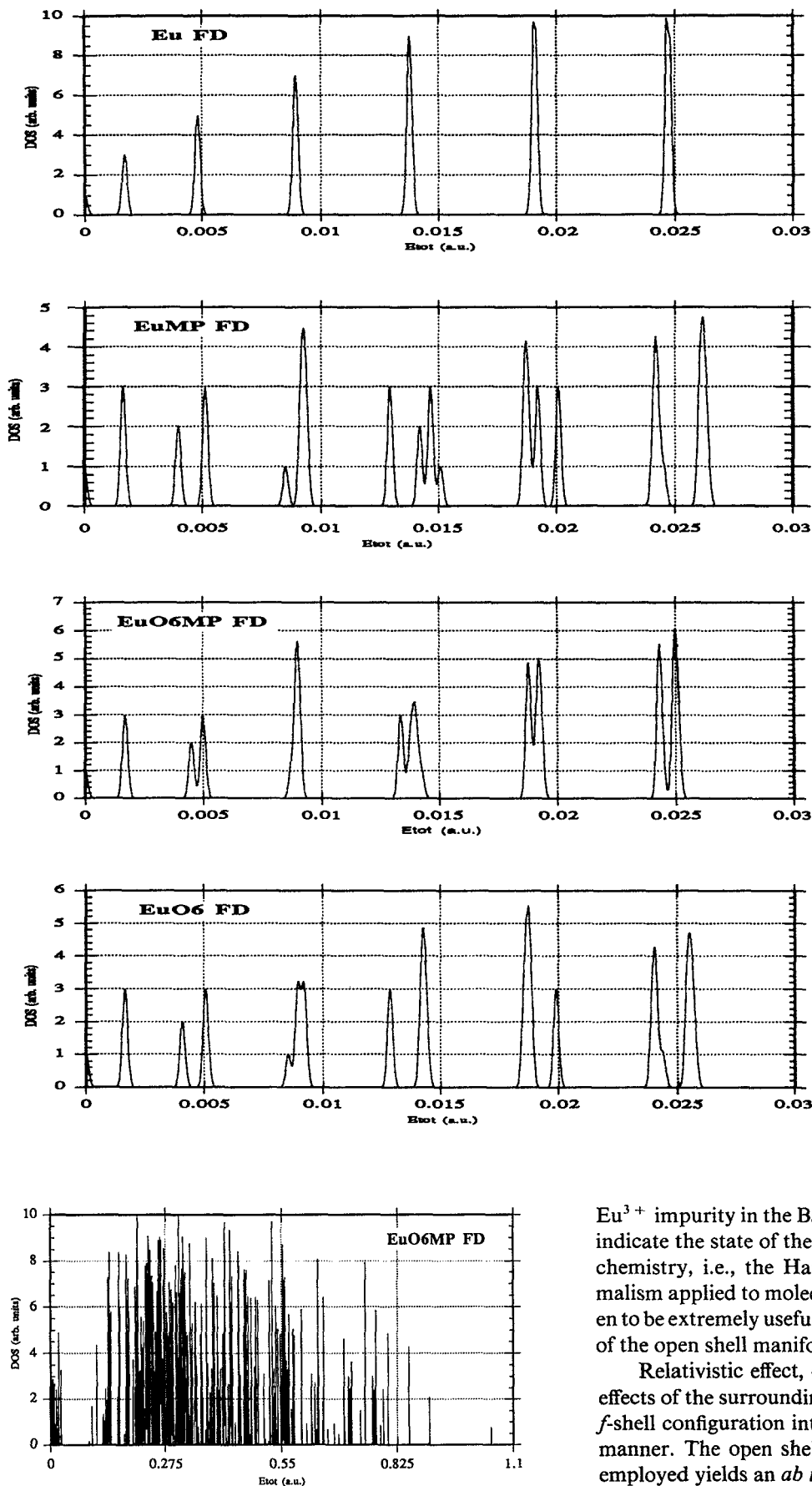


FIG. 3. Density of states from the COSCI results ( $7F$  level shown).

FIG. 4. Density of states from the  $\text{EuO}_6\text{MP FD}$  COSCI results.

$\text{Eu}^{3+}$  impurity in the  $\text{Ba}_2\text{GdNbO}_6$  crystal. The calculations indicate the state of the art in *ab initio* relativistic quantum chemistry, i.e., the Hartree–Fock–Dirac and COSCI formalism applied to molecules. The COSCI method has proven to be extremely useful in obtaining all the individual states of the open shell manifold.

Relativistic effect, except for the Breit interaction, the effects of the surroundings on the  $\text{Eu}^{3+}$  ion, and part of the  $f$ -shell configuration interaction are treated in a systematic manner. The open shell configuration interaction method employed yields an *ab initio* intermediate coupling description of the  $f^6$ -like manifold.

Our results show that the splittings due to relativistic

TABLE IX. Energies of the fluorescence transitions (in  $\text{cm}^{-1}$ ).

Transition	Eu <i>FD</i>	EuMP <i>FD</i>	EuO <sub>6</sub> MP <i>FD</i>	van Piggelen	Exp. <sup>a</sup>
$^5D_0 \rightarrow ^7F_1$	20 189	20 229	19 871	20 459	16 800
$^5D_0 \rightarrow ^7F_2 (E_g)$	19 507	19 718	19 259	19 738	16 150 <sup>b</sup>
$^5D_0 \rightarrow ^7F_2 (T_{2g})$	19 507	19 467	19 149	19 599	16 150 <sup>b</sup>

<sup>a</sup> Blasse *et al.* (Refs. 28 and 29).<sup>b</sup> Unweighted average of the six lines observed in this region. The separation between the lines is  $\sim 100 \text{ cm}^{-1}$ .

effects, which are much larger than the splittings due to the surroundings, can be calculated with an accuracy of  $\sim 10 \text{ cm}^{-1}$ . We have also found that both a pure Madelung field and a bare  $\text{EuO}_6^{9-}$  cluster significantly overestimate the splittings introduced by the surroundings. It is clear from our nonrelativistic results that the detailed form of the orbitals is important in the calculation of these splittings. Hence the preferred approach should be to use relativistic orbitals from the outset.

A significant discrepancy with experiment is found where splittings between Russell–Saunders terms are concerned (as in the luminescence spectra). This discrepancy, which is also present on the atomic level, is due to the neglect of most of the correlation effects and does not differ much from the discrepancies found in nonrelativistic calculations.

## ACKNOWLEDGMENTS

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